



1

00:00:00,000 --> 00:00:02,990

The Sun's core is the hottest part of the Sun.

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00:00:03,010 --> 00:00:06,600

But our star's temperature doesn't behave as you might expect.

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00:00:06,620 --> 00:00:09,980

The core is roughly 27 million degrees Fahrenheit

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00:00:10,000 --> 00:00:12,420

and 10 times more dense than gold.

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00:00:12,440 --> 00:00:16,930

As you move outward, the layers of the Sun become cooler and less dense.

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00:00:16,950 --> 00:00:21,320

Something unusual, however, occurs when you reach the outermost layer.

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00:00:21,340 --> 00:00:24,920

While the surface is around 10,000 degrees Fahrenheit,

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00:00:24,940 --> 00:00:29,580

the corona – the Sun's outer atmosphere — is several hundred times hotter.

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00:00:29,600 --> 00:00:34,950

That's the opposite of what happens with a fire, when it gets cooler the farther away you get.

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00:00:34,970 --> 00:00:37,520

Scientists call this the coronal heating problem

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00:00:37,540 --> 00:00:42,380

and evidence for this was first uncovered during an eclipse in the 1800s.

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00:00:42,400 --> 00:00:48,420

The corona is usually hard to see -- it's too dim to be seen next to the Sun's bright body.

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00:00:48,440 --> 00:00:50,730

But it can be seen with the naked eye,

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00:00:50,750 --> 00:00:54,190

when the moon blocks the sun during a total solar eclipse.

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00:00:54,210 --> 00:00:57,070

To understand how this mystery was discovered,

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00:00:57,090 --> 00:01:02,680

it helps to know how scientists started studying the chemical properties of materials on Earth.

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00:01:02,700 --> 00:01:06,270

In the early 1800s, instruments named spectroscopes

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00:01:06,290 --> 00:01:10,060

were invented to identify materials that emit light when heated.

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00:01:10,080 --> 00:01:16,380

Light enters the spectroscope and is filtered through a tiny hole to isolate a single area.

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00:01:16,400 --> 00:01:21,500

It then bounces off a special grating that disperses light into its different wavelengths.

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00:01:21,520 --> 00:01:24,450

While sunlight contains every wavelength,

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00:01:24,470 --> 00:01:28,090

scientists discovered that every chemical element and compound

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00:01:28,110 --> 00:01:30,180

contains a unique pattern of wavelengths

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00:01:30,200 --> 00:01:33,800

that allows scientists to determine the composition of light sources.

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00:01:33,820 --> 00:01:40,160

It wasn't long before astronomers started extracting information from the light of distant stellar objects.

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00:01:40,180 --> 00:01:48,570

In 1869, two scientists independently decided to point a spectroscope toward the corona during a total solar ec

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00:01:48,590 --> 00:01:53,150

As the Sun's light disappeared, the pattern of wavelengths changed.

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00:01:53,170 --> 00:01:55,580

They saw something they had never seen before.

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00:01:55,600 --> 00:02:00,930

A bright green line that did not relate to any element found on Earth.

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00:02:00,950 --> 00:02:04,180

For a short while, scientists named it 'coronium'.

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00:02:04,200 --> 00:02:08,610

It wasn't until 70 years later that a Swedish scientist discovered

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00:02:08,630 --> 00:02:14,110

that these lines were the result of elements such as iron being stripped of its electrons.

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00:02:14,130 --> 00:02:18,440

Every element has a specific number of electrons surrounding the nucleus.

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00:02:18,460 --> 00:02:23,050

As each electron is removed, more energy is needed to remove the next one.

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00:02:23,070 --> 00:02:28,830

The green line shows that iron has been stripped of 13 of its 26 electrons –

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00:02:28,850 --> 00:02:35,870

– indicating that the corona needed to be millions of degrees -- counterintuitively far hotter than the Sun's surfa

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00:02:35,890 --> 00:02:39,670

Scientists have since proposed a variety of theories

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00:02:39,690 --> 00:02:43,560

for what mechanisms could be adding that extra heat into the atmosphere.

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00:02:43,580 --> 00:02:50,320

One theory suggests that small waves in the Sun's surface pushes particles and heat into the atmosphere

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00:02:50,340 --> 00:02:53,210

- a bit like how ocean waves push surfers.

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00:02:53,230 --> 00:03:00,090

Another theory suggests small bomb-like explosions from the realignment of the Sun's magnetic field create heat

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00:03:00,110 --> 00:03:02,690

Many scientists think it may be a mix of both.